# An Update on the Radio Source WISE J071634.59-190039.2

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Image courtesy NRAO/AUI

## W0716 and FRB150418 were claimed to be linked.

The FRB and variable radio source were discovered by the SUPERB team using Parkes and ATCA (Keane+ 2016).



Further follow-up revealed W0716 is a persistently variable source.

### We have new data: FRB121102 was localized!

Worth revisiting the potential association between W0716 and FRB150418 in light of this.

The attention-getting piece: FRB121102 is associated with a persistent, variable, compact radio source.

Standard caveats about extrapolating from a sample size of one apply.



5 h 32 min 12 s 5 h 32 min 0 s 5 h 31 min 48 s 5 h 31 min 36 s Right ascension



Chatterjee+ (2017)

## The two galaxies do not look very similar.

W0716 is a passive elliptical. The 121102 host is a low-metallicity, star-forming dwarf.

Property	W0716	Host 121102	Unit
Redshift	0.49	0.19	
M <sub>r'</sub>	-21.6	-17.0	mag
Diameter	~22	≲4	kpc
M*	10 <sup>11</sup>	10 <sup>7.7</sup>	${\sf M}_{\odot}$
SFR	≲0.2	0.4	M <sub>☉</sub> /yr
L <sub>X</sub>	≲10 <sup>43.8</sup>	≲10 <sup>41.5</sup>	erg/s

#### The two radio sources have similar luminosities.



Derived from Vedantham+ (2016), Chatterjee+ (2017)

## W0716 is much more variable.

The shape of W0716's spectrum ( $\rightarrow$ ) varies much more dramatically than that of the 121102 host ( $\downarrow$ ).



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Derived from PKGW & Berger (2016), Chatterjee+ (2017)

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## EVN data give more insight into the radio sources.

Property	W0716	Host 121102	Unit
$L_{v}$	10 <sup>30</sup>	10 <sup>29.5</sup>	erg/s/Hz
Τ <sub>Β</sub>	≳10 <sup>8.5</sup>	≥10 <sup>7.7</sup>	K
Variability	High	Moderate	
Size	≲1.2	≲0.7	рс
Spectrum	Flat?	Flat + turnover	
Host offset	?	~700?	рс

If the variability is due to scintillation, there is not much to distinguish the two sources.

#### <u>Here's a summary.</u>

- The two host galaxies are dissimilar.
- The clearest difference between the two radio sources is likely due to an *extrinsic* factor.
- It remains true that W0716 looks a whole lot like an AGN.
- More data could potentially strengthen the circumstantial case for associating W0716 and FRB150418 ...
- ... but direct localizations are vastly more preferable!

*Thanks for your attention!* 

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